

RADC-TR-77-332 IN-HOUSE REPORT SEPTEMBER 1977



JAN 5 1978

Radiation Effects on AlGaAs/GaAs Solar Cells Using 0.9 - 3.0 MeV Protons and 1.0 - 1.4 MeV Electrons

L.F. LOWE J.R. CAPPELLI L.W. JAMES R.L. MOON

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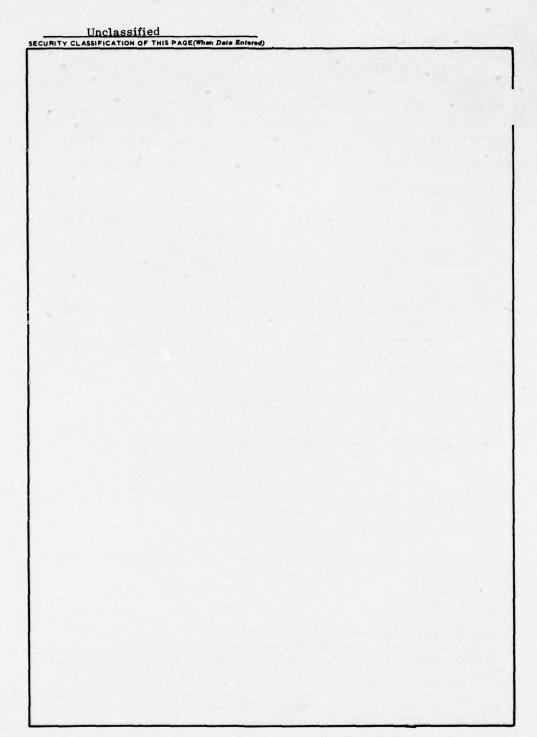
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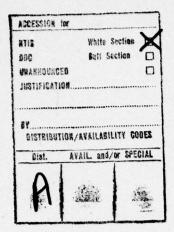
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BEFORE COMPLETING FORM REPORT DOCUMENTATION PAGE RADC-TR-77-332 rept TITLE (and Subtitle) RADIATION EFFECTS ON AlGaAs/GaAs SOLAR CELLS USING Ø. 9 - 3.0 MeV PROTONS AND 1.0 - 1.4 MeV ELECTRONS 1 Jan, -31 Dec 20076 TITHOR(a) 8. CONTRACT OR GRANT NUMBER(s) L. F. Lowe, J. R. Cappelli, L. W. James, R. L. Moon PERFORMING ORGANIZATION NAME AND ADDRESS Deputy for Electronic Technology (RADC/ESR) Hanscom AFB Massachusetts 01731 46002001 1. CONTROLLING OFFICE NAME AND ADDRESS 2. REPORT DAT Deputy for Electronic Technology (RADC/ESR September 1977 Hanscom AFB 33 Massachusetts 01731 15. SECURITY CLASS. (of this report) MONITORING AGENCY NAME & ADDRESS(if different from Controlling Office) Unclassified 15a. DECLASSIFICATION/DOWNGRADING 6. DISTRIBUTION STATEMENT (of this Report) Approved for public release; distribution unlimited. 17. DISTRIBUTION STATEMENT (of the abstract entered in Block 20, if different from Report) 18. SUPPLEMENTARY NOTES *Varian Associates, Palo Alto, California 19. KEY WORDS (Continue on reverse side if necessary and identify by block number) (1) Solar cells (2) AlGaAs Radiation effects (4) Space environment 26. ABSTRACT (Continue on reverse side if necessary and identify by block number) Aluminum gallium arsenide solar cells were irradiated with 1.0 and 1.4 MeV electrons, and with 0.9 and 3.0 MeV protons to determine radiation sensitivity. Electron fluences ranged from 1×10^{12} to 3×10^{16} electrons cm², and proton fluences from 5×10^{16} to 2.7×10^{12} cm². A solar simulator and a tungsten lamp were used to evaluate changes in the current-voltage characteristic curves. In most cases, AlGaAs solar cells showed a greater resistance to radiation than silicon cells. -10 to the 12th power DD 1 JAN 73 1473 EDITION OF 1 NOV 65 IS OBSOLETE Unclassified SECURITY CLASSIFICATION OF THIS PAGE (When Date Entered) 10 to the 14th power 10 to the 16th power 10 to the 10th power /sgcm

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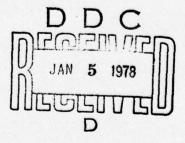


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Preface

The authors thank Roger Little and John Menucci of Simulation Physics, Inc. for their assistance in the solar simulator and tungsten lamp measurements.



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Radiation Effects on A1GaAs/GaAs Solar Cells Using 0.9 – 3.0 MeV Protons and 1.0 – 1.4 MeV Electrons

1. INTRODUCTION

In order to evaluate the applicability of AlGaAs for use as a solar cell material, information is needed on its sensitivity to the radiation that they would be exposed to in space. To this end AlGaAs solar cells were exposed to 1.0 and 1.4 MeV electrons, and 0.9 and 3.0 MeV protons. The electron fluences ranged from 1×10^{14} to 3×10^{16} electrons cm⁻², and the proton fluences ranged from 5×10^{10} to 2.7×10^{12} protons cm⁻².

2. EXPERIMENTAL

Cells were fabricated by conventional LPE processes, where the p-type AlGaAs (AlAs - 0.80) was grown on an n-type GaAs(111)B substrate. During growth, the p-dopant diffused into the n-layer to form a p-n junction approximately 0.3 to 0.8 micron deep in the GaAs. Doping densities of the p-AlGaAs + n-GaAs substrate were 3×10^{18} cm⁻³ and 8×10^{17} cm⁻³, respectively. A p-type GaAs layer was grown on top of the AlGaAs in order to facilitate contacting, and later removed by a selective etch using the contact grid as a mask. Cell area was 1.613 cm². These cells were not optimized for air mass zero (AMO) performance,

(Received for publication 29 September 1977)

since the AlGaAs were generally too thick, reducing the blue response. SiO_2 deposited to a thickness of 750 Å was used as an antireflection coating. AMO efficiencies were 10-12 percent. Proton irradiation experiments were conducted at the Van de Graaff facility at RADC/ET, Massachusetts. After each irradiation to a given fluence, the I-V curve was measured using an AMO solar simulator and a tungsten lamp as light source. Proton energies of 0.9 and 3 MeV were used, up to maximum fluences of $5 \times 10^{11} \text{ p/cm}^2$ and $3 \times 10^{12} \text{ p/cm}^2$, respectively.

Electron bombardment was carried out at the Dynamitron accelerator at RADC/ET, Massachusetts. Electron energies of 1 and 1.4 MeV were used up to maximum fluence of $3\times10^{16}~\rm e/cm^2$ and $3\times10^{15}~\rm e/cm^2$, respectively. AMO data were taken using Spectrolab solar simulators at these locations. All measurements were taken at 25° C.

After initial characterization, all samples were evaluated for radiation induced changes in their current-voltage (i-v) characteristic curves using both a solar simulator at AMO and also a tungsten lamp. After each tungsten run, the tungsten intensity was adjusted so as to bring the short circuit current back to what it was before any irradiation (I_{SCO}). In addition after the final irradiation, the spectral response was determined for each cell. The individual irradiation procedures were as follows.

3. ELECTRONS

The electron irradiations were made using a 1.5 MeV Dynamitron electron accelerator. Two samples each were mounted on either side of the aperture of a Faraday cup so that all fluences were measured directly. The cells were open circuited during the irradiation and the exposures were carried out at room temperature in air. The electron flux was approximately 10¹² electrons cm⁻² s⁻¹. Two energies were used, 1.0 and 1.4 MeV.

4. PROTONS

A 3.0 MeV Van de Graaff accelerator was used for the proton exposures. Because of the short range of protons these experiments were carried out in vacuum, one cell at a time. Each cell was mounted in the center of a Faraday cup which was part of the vacuum system. A collimator, electrically isolated from the Faraday cup, defined the exposure area. The Van de Graaff is equipped with a beam sweeping device which leads to an exposure uniformity of better than 10 percent. Proton flux levels were 10¹⁰ protons cm⁻² s⁻¹, and runs were made open circuited at 0.9 and 3.0 MeV.

5. RESULTS

Figures 1-8 give the individual I-V curves as a function of fluence using the solar simulator at AMO. Figures 9-16 show tungsten lamp data for the same exposures. Figures 17-21 illustrate the effect of increasing the tungsten lamp intensity so as to bring the cell output current back to its preirradiation value (the I_{sco} case). Contact problems were encountered with 3 cells. Pertinent data from Figures 1-16 are given in Tables 1-8, including the incomplete data for the three cells with contact problems. The output power data is shown in Figures 22-26. After two weeks, the cells were checked for annealing. None was observed. Spectral response data are shown in Figures 27-30.

6. SUMMARY AND DISCUSSION

Solar cells of p-AlGaAs/n-GaAs were subjected to 0.9 and 3 MeV proton radiation, and 1 and 1.4 MeV electron radiation. Values of critical fluence \emptyset_c at which the output power was reduced by 25 percent were determined. As expected protons with energies of 0.9 MeV caused the most severe degradation, and a $\emptyset_c = 4 \times 10^{10} \text{ p/cm}^2$ was determined. Fortunately because of the short range of low energy protons, this type of radiation can be protected against by cover slips. For 3 MeV protons, $\emptyset_c = 5 \times 10^{11} \text{ p/cm}^2$.

Critical fluences for 1.4 MeV electrons were $7-9\times10^{14}~\rm e/cm^2$. A total of five cells were irradiated at electron energies of 1 MeV. $\rm \phi_c$ varied between $7\times10^{14}~\rm and 7\times10^{15}~\rm e/cm^2$. At fluences $\rm <10^{15}~\rm e/cm^2$, the relative cell parameters followed the characteristic radiation equation. At fluences greater than this the relative parameters varied linearly with $\rm ln0$, as did the relative diffusion length.

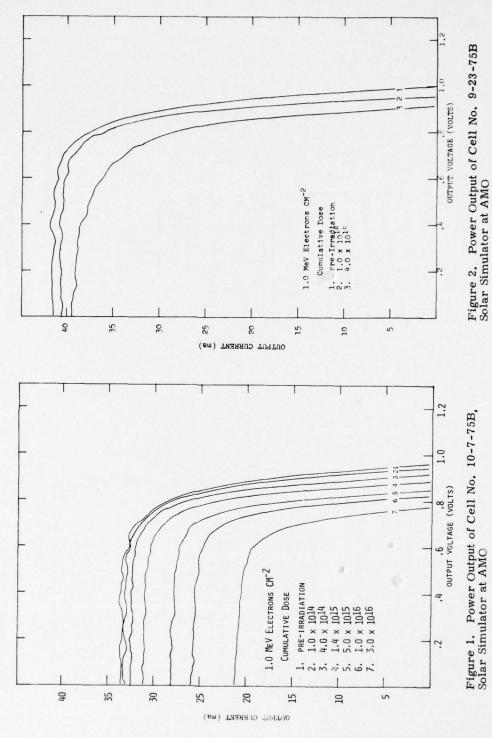
These results show that the radiation resistance of AlGaAs/GaAs solar cells is equal to or better than that observed in conventional and violet Si cells for all the radiation fluxes investigated here, except 0.9 MeV protons. In the case of 1 MeV electrons, the value of $\emptyset_{\rm c}=7\times10^{15}~{\rm e/cm}^2$ is believed to be the highest observed for conventional cells. The spread in $\emptyset_{\rm c}$ we observed is thought to be caused by variations in junction depth, which can be optimized. So even for 1 MeV electrons, cells of AlGaAs/GaAs are superior to those of Si.

Computer calculations show that the variation in I_{sc} as a function of irradiation level is sensitive to junction depth. Thus a shallow junction solar cell should be more radiation-resistant than the experimental data shown here, since I_{sc} does not change as rapidly with decreasing diffusion length as in a deep junction device.

Calculations also demonstrated that the AMO efficiency increases with AlAs concentration in the AlGaAs for a given thickness. This will allow operation in

concentrated sunlight in the space environment because the sheet resistance of the AlGaAs can be reduced without substantially degrading performance through optical absorption of high energy protons in the contact layer.

Combining the radiation resistance characteristics with the ability to obtain AMO efficiencies of 16 percent, and potentially 18 percent, makes the AlGaAs/GaAs solar cells attractive for space applications.



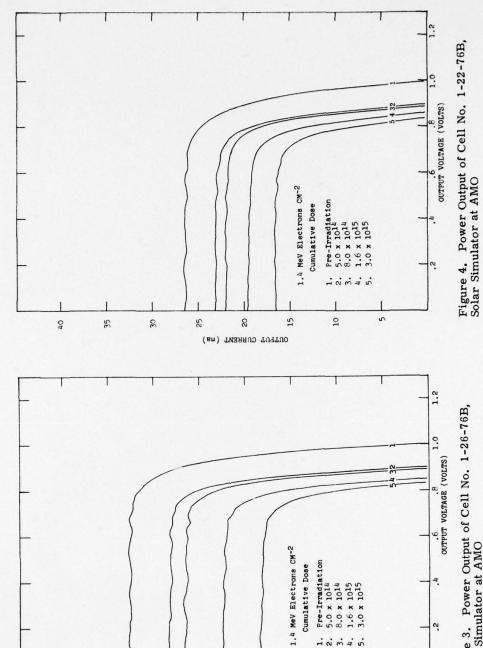
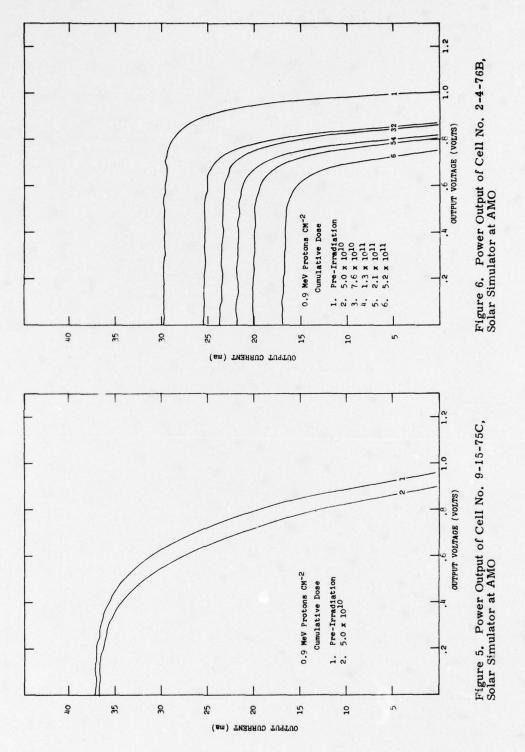


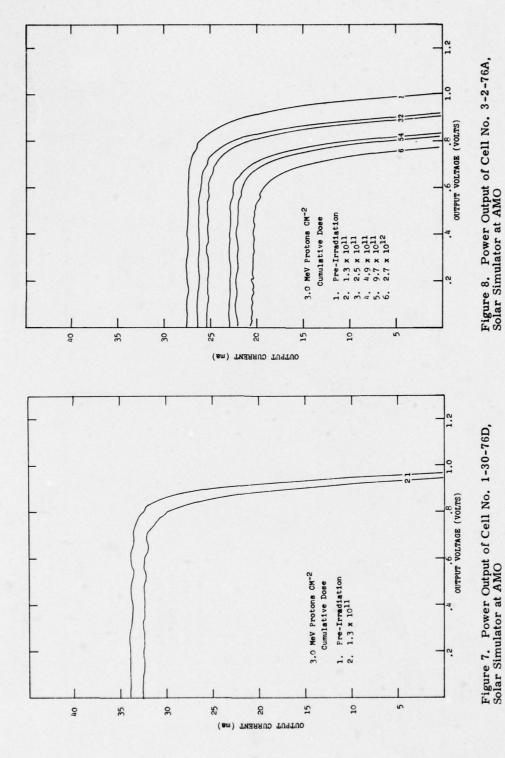
Figure 3. Power Output of Cell No. 1-26-76B, Solar Simulator at AMO

OUTPUT CURRENT (ma)

Pre-Irradiation 5.0 x 101⁴ 8.0 x 101⁴ 1.6 x 1015 3.0 x 1015

Cumulative Dose





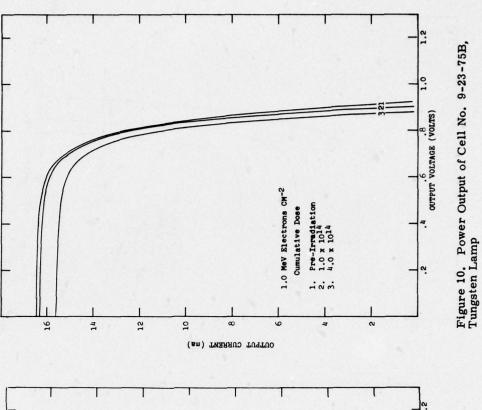


Figure 9. Power Output of Cell No. 10-7-75B, Tungsten Lamp

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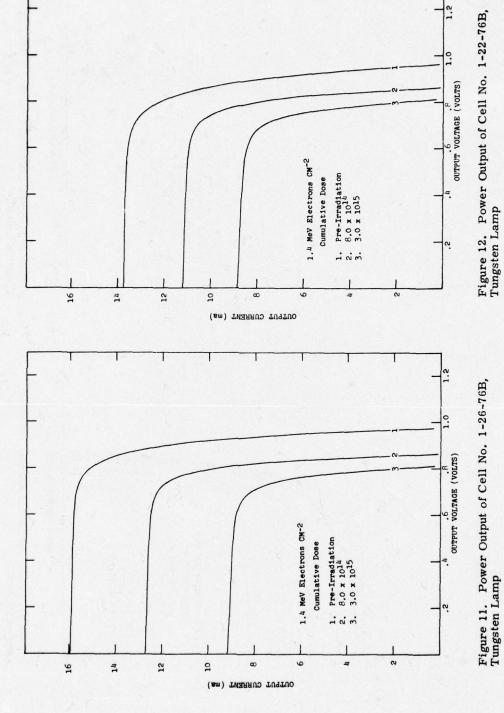
16

14

12

10

1.0 MeV Electrons CM-2 Cumulative Dose



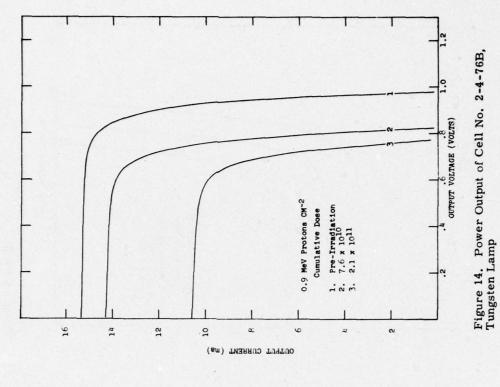


Figure 13. Power Output of Celi No. 9-15-75C, Tungsten Lamp

OUTPUT VOLTAGE (VOLTS)

8

OUTPUT CURRENT (ma)

16

14

12

10

Cumulative Dose

1. Pre-Irradiation
2. 9.6 x 10¹⁰

N

9

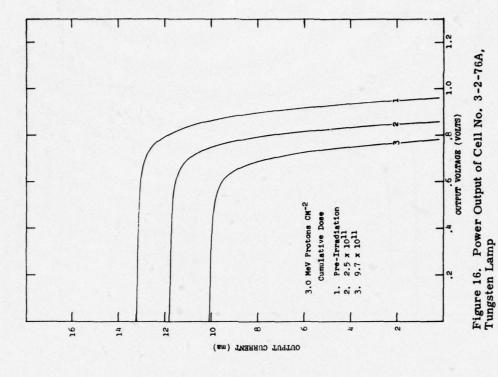


Figure 15. Power Output of Cell No. 1-30-76D, Tungsten Lamp

OUTPUT VOLTAGE (VOLTS)

OUTPUT CURRENT (ma)

10

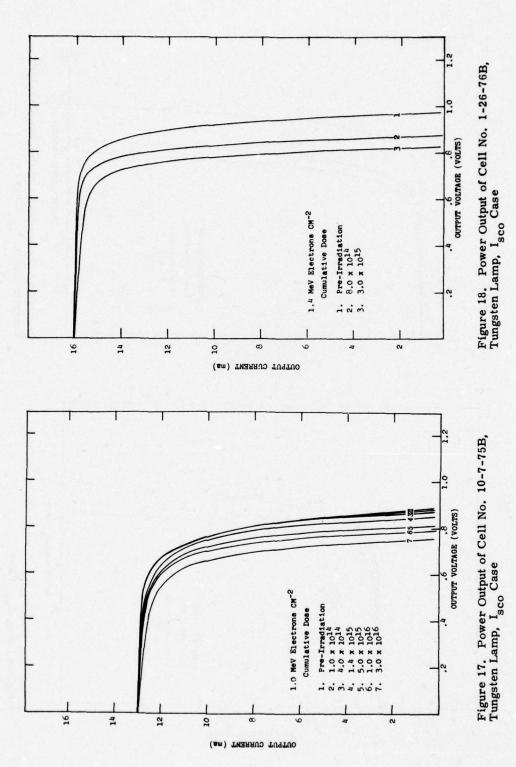
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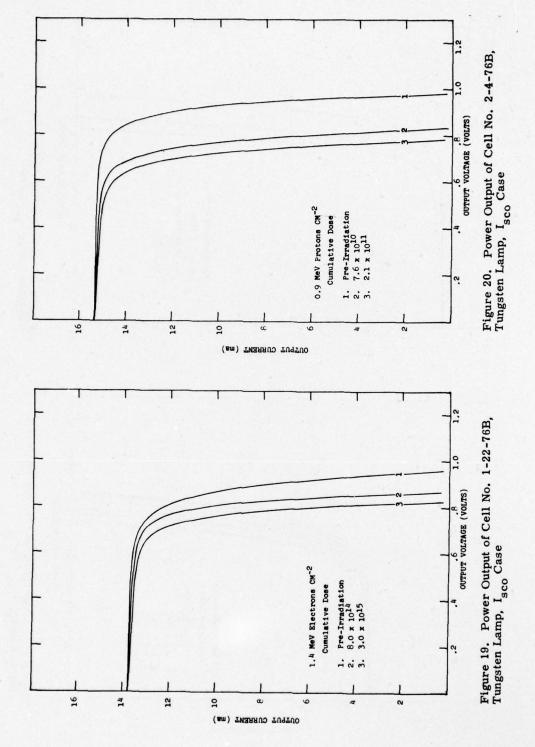
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16

3.0 MeV Protons CM-2

1. Pre-Irradiation 2. 2.6 x 10¹¹ Cumulative Dose





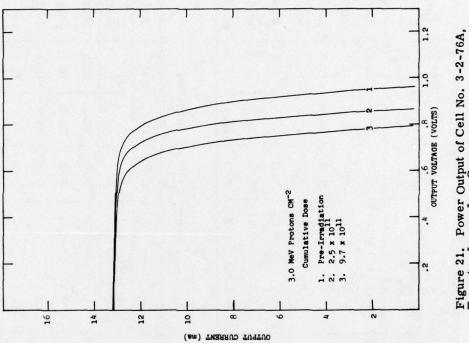


Figure 21. Power Output of Cell No. 3-2-76A, Tungsten Lamp, $\rm I_{sco}$ Case

Table 1. Cell No. 10-7-75B

		I _{sc} (mA)	nA)	Voc (volts)	volts)	P _{ep} (mw)	mw)	Pmax (mw)	(mw)	Fill Factor	actor
	e cm-2	Solar Simulator	Tungsten Lamp	s s	LT	w w	LT	w w	LI	w w	I
	0	33, 5	13.0	96.0	0.89	32.2	11.6	22.8	8.1	0,708	0, 698
-	1.0×10^{14}	33.0	13.0	0.94	0.88	31.0	11.4	22.6	8.1	0.729	0,711
4	4.0×10^{14}	32.5	12.6	0.91	0.87	29.6	11.0	22.0	8.7	0.743	0,709
_	1.4×10^{15}	31.1	11.9	0.88	0.85	27.4	10.1	20.3	7.3	0.741	0.723
5,	5.0×10^{15}	28.0	10.7	0.85	0.81	23.8	8.7	17.9	6.4	0.752	0,736
	1.0×10^{16}	26.0	9.7	0.82	0.78	21.3	7.6	15.7	5.4	0.737	0,711
3,	3.0×10^{16}	21.3	8.2	0.77	0.74	16.4	6.1	11.6	4.3	0.707	0,705

Table 2. Cell No. 9-23-75B

	$_{\rm sc}^{\rm I}$	I _{sc} (mA)	V _{oc} (Voc (volts)	P _{ep} (mw)	(mm)	P _{max} (mw)	(mm)	Fill Factor	actor	
e_ cm_2	Solar Simulator	Tungsten Lamp	ωω	T	ωω	T	ωω	T	S S	LT	
	41.5	16.5	0.99	0.92	41.1	15.2	30.0	30.0 10.7	0.730	0,704	
1.0×10^{14}	40.5	16.3	96.0	06.0	38.9	14.7	28.9	10.7	0.743	0.728	
4.0×10^{14}	39, 5	15.6	0.91	0.88	35.9	13.7	24.1	10.0	0.671	0.730	
1.4×10^{15}	35.5	14.5	1	0.82	;	11.9	:	1	!	:	
2.5×10^{16}	19.5	:	0.78	:	15.4	;	11.3	;	0.734	;	

Table 3. Cell No. 1-26-76B

I _{SC} (mA) V _{oc} (volts) Pep (mw) Solar Tungsten S T S T Simulator Lamp S L S L 32.6 15.9 1.0 0.97 32.6 15.4 28.5 0.90 25.7 26.6 12.7 0.89 0.86 23.7 10.9 22.6 0.85 19.2											
Solar Tungsten S T S T Simulator Lamp S L S L 32.6 15.9 1.0 0.97 32.6 15.4 28.5 0.90 25.7 26.6 12.7 0.89 0.86 23.7 10.9 22.6 0.85 19.2		m) ^{os} l	(A)	V _{oc} ((volts)	P _{ep} (mw)	Pmax (mw)	(mm)	Fill Factor	actor
32.6 15.9 1.0 0.97 32.6 15.4 28.5 0.90 25.7 26.6 12.7 0.89 0.86 23.7 10.9 22.6 0.85 19.2	e cm-2	Solar Simulator	Tungsten	w w	FJ	w w	FJ	w w	LI	SS	I
28.5 0.90 25.7 26.6 12.7 0.89 0.86 23.7 10.9 22.6 0.85 19.2	0	32.6	15.9	1.0	0.97	-	15.4	25.8	12.1	0, 791	0,786
26.6 12.7 0.89 0.86 23.7 10.9 22.6 0.85 19.2	5.0×10^{14}	28.5	1	0.90	+	25.7	1	20.5	1	0.798	;
22.6 0.85 19.2	8.0×10^{14}	26.6	12.7	0.89	0.86	23.7	10.9	19.0	8.7	0.802	0.798
6 0 0	1.6×10^{15}	22.6	1	0.85	1	19.2	:	15.1	;	0.786	ì
18.4 9.2 0.83 0.81 15.3 (.3	3.0×10^{15}	18.4	9.2	0.83	0.81	15.3	7.5	11.9	5.7	0.778	092.0

Table 4. Cell No. 1-22-76B

	I _{sc} (mA)	(Ar	Voc (Voc (volts)	Pep	Pep (mw)	Pmax (mw)	(mw)	Fill Factor	actor
e cm -2	Solar	Tungsten Lamp	w w	FJ	w w	LI	S S	T	S	$rac{ ext{T}}{ ext{L}}$
0	26.5	13.8	0.99	96.0	26.2	13.2	19.8	8.6	0.756	0.742
5.0×10^{14}	23.0	1	0.89	1	20.5	1	16.1	1	0.785	;
8.0×10^{14}	22.0	11.2	0.89	0.86	19.6	9.6	15.4	7.5	0.786	0.781
1.6×10^{15}	19.5	1	0.86	1	16.8	1	13.0	1	0.774	1
3.0×10^{15}	16.7	8.8	0.83	0.81	13.9	7.1	10.7	5.4	0.770	0.761

Table 5. Cell No. 9-15-75C

П				
Fill Factor	T L	0,644	1	0.633
Fill F	SS	0,534	0.505	1
(mm)	T L	8.5	1	6.9
P _{max} (mw)	SS	18.9	16.5	:
mw)	T L	13.2	1	10.9
Pep (mw)	s s	35.4	32.7	1
Voc (volts)	T	0.91	1	0.81
V _{oc} (S S	0.95	0.89	;
A)	Tungsten Lamp	14.5	1	13.5
I _{sc} (mA)	Solar Simulator	37.3	36.7	33,5
	e cm-2	0	5.0×10^{10}	9.6×10^{10}

Table 6. Cell No. 2-4-76B

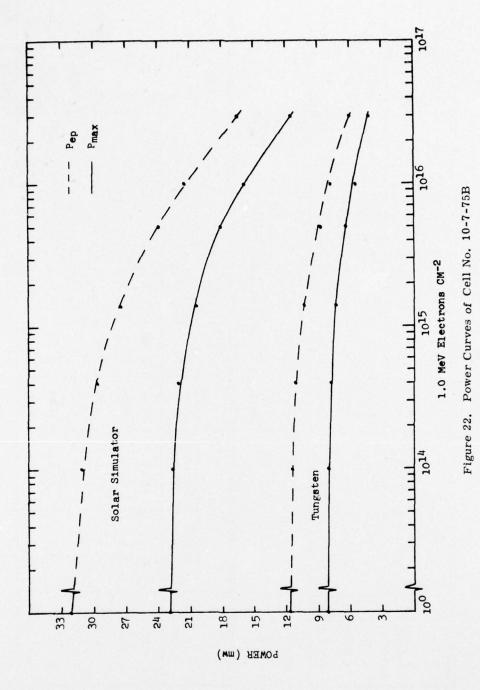
	I _{sc} (mA)	(A)	Voc (volts)	volts)	Pep (mw)	mw)	Pmax (mw)	(mm)	Fill Factor	actor
e cm-2	Solar Simulator	Tungsten Lamp	S S	T	S S	FI	w w	LI	w w	LI
0	29.7	15.3	1.0	96.0	29.7	15.0	23.5	11.8	0, 791	0.787
5.0×10^{10}	25.5	-	0.87	1	22.2	1	17.2	;	0.775	;
7.6×10^{10}	23.7	12.0	0.86	0.81	20.4	9.7	15.5	7.4	092.0	0,763
1.3×10^{11} -	21.9	1	0.82	1	18.0	1	13.4	1	0.744	
2.1×10^{11}	20.1	10.6	08.0	0.77	16.1	8.2	12.1	0.9	0.752	0.732
5.2×10^{11}	17.0	ŀ	0.75	1	12.8	1	9.3	1	0.727	1

Table 7. Cell No. 1-30-76D

	I _{sc} (mA)	nA)	Voc (volts)	volts)	Pep (mw)	nw)	Pmax (mw)	(mm)	Fill Factor	ctor
e_ cm_2	Solar Simulator	Tungsten Lamp	S S	T	s s	T	S S	T	ωω	T L
0	33,8	14.0	76.0	0,93	32.8	13.0	26.6	10.5	0.811	0,808
1.3×10^{11}	32, 5	1	0.94	1	30.6	1	23.8	1	0.778	;
2.6×10^{11}	1	13.1	;	0.89	1	11.7	1	8.9	1	0.761

Table 8. Cell No. 3-2-76A

	I _{sc} (mA)	A)	V _{oc} (Voc (volts)	Pep (mw)	(mm)	P _{max} (mw)	(mm)	Fill Factor	actor
e cm-2	Solar Simulator	Tungsten Lamp	w w	I	ωω	LT	s s	LI	w w	L
0	27.5	13.2	1.0	96.0	27.5	12.7	21.0	9.6	0.764	0.756
1.3×10^{11}	26.4	1	0.92	+	24.3	1	18.4	1	0.757	:
2.5×10^{11}	25.5	11.8	0.91	0.86	23.2	10.1	17.5	7.7	0,754	0,762
4.9×10^{11}	23.0	-	0,83	;	19.1	!	14.4	1	0.754	;
9.7×10^{11}	22.1	10.1	0.82	0.78	18.1	6.7	13.6	5.8	0,752	0.734
2.7×10^{12}	20.5	-	0.77	1	15.8	-	11.7	1	0,741	;



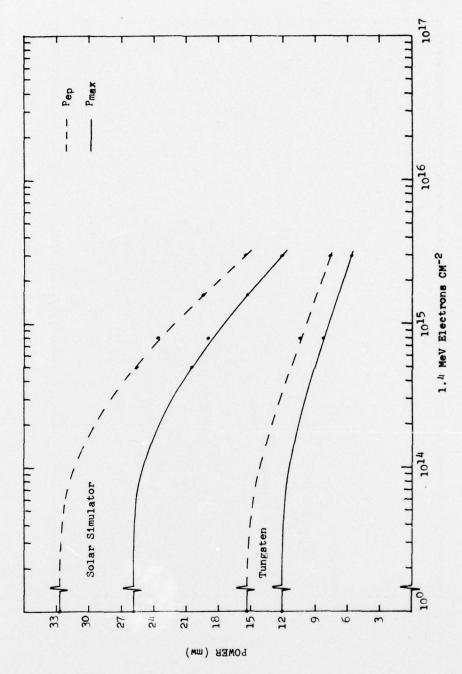


Figure 23. Power Curves of Cell No. 1-26-76B

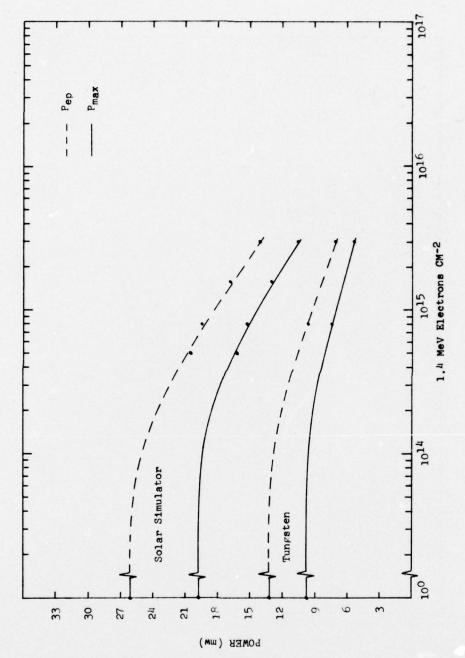


Figure 24. Power Curves of Cell No. 1-22-76B

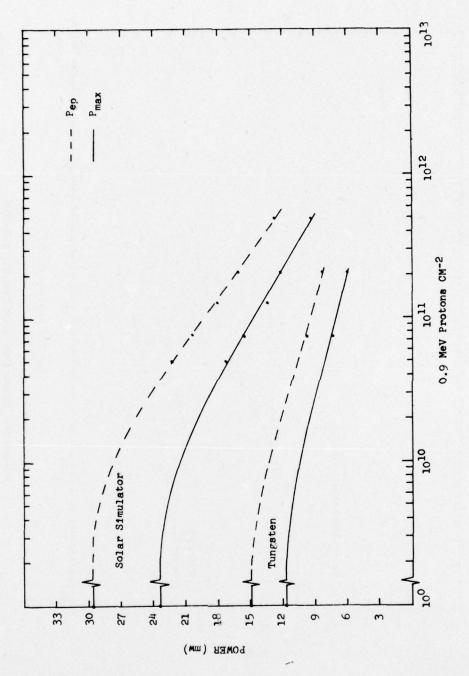


Figure 25. Power Curves of Cell No. 2-4-76B

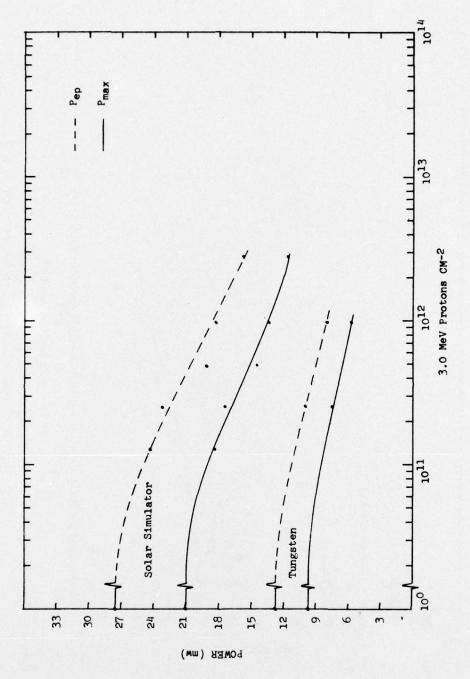


Figure 26. Power Curves of Cell No. 3-2-76A

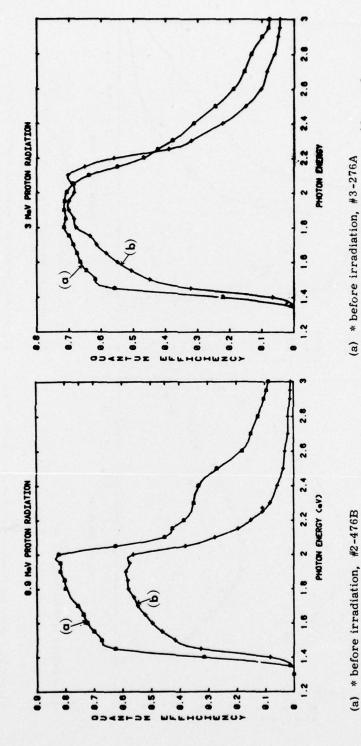


Figure 27. Spectral Response Curve for a Solar Cell Irradiated With 0.9-MeV Proton

(b) + after irradiation to 0 = 2.7 $\times\,10^{12}~\mathrm{p/cm}^2$

(b) + after irradiation to 0 = 9.6 $\times\,10^{10}~\mathrm{p/cm}^2$

(a) * before irradiation, #2-476B

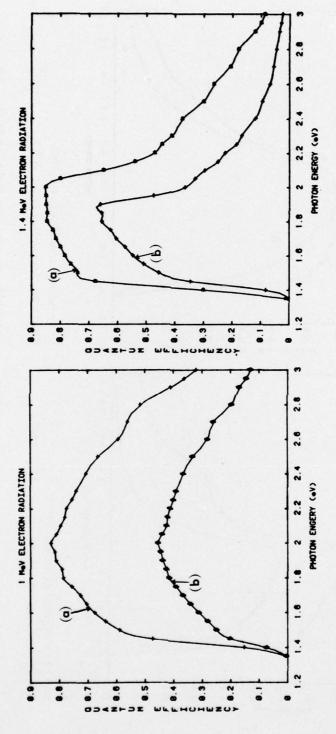


Figure 30. Spectral Response Curve for a Solar Cell Irradiated With 1.4-MeV Electrons Figure 29. Spectral Response Curve for a Solar Cell Irradiated With 1-MeV Electron

(b) + after irradiation to \emptyset = 3×10^{15} e/cm²

(b) + after irradiation to 0 = 3×10^{16} e/cm²

(a) * before irradiation, #9-2375B

(a) * before irradiation, #1-2276B

METRIC SYSTEM

METRIC SYSTEM				
BASE UNITS:				
Quantity	Unit	SI Symbol	Formula	
length	metre	m		
mass	kilogram	kg	***	
time	second	5		
electric current	ampere	٨		
thermodynamic temperature	kelvin	K	***	
amount of substance	mole	mol	***	
luminous intensity	candela	cd	***	
SUPPLEMENTARY UNITS:				
plane angle	radian	rad		
solid angle	steradian	ST		
DERIVED UNITS:				
Acceleration	metre per second squared		m/s	
activity (of a radioactive source)	disintegration per second		(disintegration)/s	
angular acceleration	radian per second squared		rad/s	
angular acceleration	radian per second		rad/s	
	square metre		m	
area	kilogram per cubic metre		kg/m	
density	farad	F	A-s/V	
electric capacitance	siemens	S	AN	
electrical conductance	volt per metre		V/m	
electric field strength	henry	н	V-s/A	
electric inductance		Ÿ	W/A	
electric potential difference	volt ohm		V/A	
electric resistance	volt	V	W/A	
electromotive force			N-m	
energy.	joule		1/K	
entropy	joule per kelvin	Ň	kg-m/s	
force	newton	Hz	(cycle)/s	
frequency	hertz	lx	lm/m	
illuminance	lux		cd/m	
luminance	candela per square metre	lm	cd-sr	
luminous flux	lumen	ım	A/m	
magnetic field strength	ampere per metre	Wb	V-s	
magnetic flux	weber	T	Wb/m	
magnetic flux density	tesla		***************************************	
magnetomotive force	ampere	A w]/s	
power	watt	VV Pa	N/m	
pressure	pascal	C	A·s	
quantity of electricity	coulomb		N·m	
quantity of heat	joule		W/sr	
radiant intensity	watt per steradian	***		
specific heat	joule per kilogram-kelvin	<u> </u>	J/kg-K N/m	
stress	pascal	Pa	N/m W/m-K	
thermal conductivity	watt per metre-kelvin	***		
velocity	metre per second	***	m/s	
viscosity, dynamic	pascal-second		Pa·s	
viscosity, kinematic	square metre per second	***	m/s	
voltage	volt	v	W/A	
volume	cubic metre		m	
wavenumber	reciprocal metre	***	(wave)/m	
work	joule	1	N-m	
SI PREFIXES:				
Multiplication Factors		Prefix	SI Symbol	
1,000,000	000 000 = 1012	tera	Т	
	000 000 = 109	giga	G	
	000 000 = 10 ⁶	mega	M	
1 000 000 = 10		kilo	k	

Multiplication Factors	Prefix	SI Symbol
1 000 000 000 000 = 1012	tera	т
1 000 000 000 = 104	giga	G
1 000 000 = 106	mega	M
1 000 = 103	kilo	k
100 = 102	hecto*	h
10 = 10'	deka*	de
0.1 = 10-1	deci*	d
$0.01 = 10^{-2}$	centi*	C
0.001 = 10~3	milli	m
0.000 001 = 10~4	micro	μ
0.000 000 001 = 10-9	nano	n
0.000 000 000 001 = 10-12	pico	P
0.000 000 000 000 001 = 10-15	femto	1
0.000 000 000 000 000 001 = 10-18	atto	•

^{*} To be avoided where possible

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are communications, el'

surveillance of grow data collection ar ionospheric prophysics and e'

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